

Gravitation Markscheme

1. A 1
2. C [1]
3. B [1]
4. C+D [1]
5. C [1]
6. B [1]
7. A [1]
8. B [1]
9. C [1]
10. B [1]
11. A [1]
12. (a) the force exerted per unit mass;
on a point (small) mass; 2
- (b) (i) use of $g = \frac{F}{m}$ and $F = G \frac{Mm}{R^2}$;
combine to get $g = G \frac{M}{R^2}$; 2
- (ii) $M = \frac{gR^2}{G}$;
substitute to get $M = 1.9 \times 10^{27}$ kg; 2
- 13. Kepler's third law**
- (a) (i) the orbits are elliptical / not circular;
Do not accept "because R changes". 1
- (ii) gravity / gravitational; 1
- (iii) $F = G \frac{M_1 m}{R^2}$; 1
- Same symbols as in question must be used to receive the mark.*

1

- (iv) let v = the speed of planet, the acceleration is then = $\frac{v^2}{R}$
from Newton 2, $F = G \frac{M_s m}{R^2} = \frac{mv^2}{R}$; *{ Some reference to Newton 2
is required to receive this mark.*

$$v = \frac{2\pi R}{T};$$

$$\text{therefore, } = G \frac{M_s m}{R^2} = \frac{4\pi^2 m R^2}{T^2};$$

$$\text{therefore, } T^2 = \frac{4\pi^2}{GM_s} R^3; \quad 4$$

$$\text{so } K = \frac{4\pi^2}{GM_s}$$

Be aware of the many simple variants of this eg

$$\text{using } a = \omega^2 R = \frac{4\pi^2 R}{T^2}.$$

- (b) (i) recognize that gravitational field strength = $\frac{v^2}{R} = \frac{4\pi^2 R}{T^2}$;

$$= \frac{40 \times 1.1 \times 10^9}{(6.2 \times 10^5)^2};$$

$$= 0.1 \text{ N kg}^{-1} \quad 2$$

- (ii) $T^2 = \frac{4\pi^2}{GM_s} R^3$ therefore, $M_{\text{Jup}} = \frac{4\pi^2 R^3}{GT^2}$;

$$M_{\text{Jup}} = \frac{4\pi^2 \times (1.1)^3 \times 10^{27}}{6.7 \times 10^{-11} \times (6.2 \times 10^5)^2};$$

$$\approx 2.0 \times 10^{27} \text{ kg};$$

Or

$$0.1 = \frac{GM}{R^2};$$

$$M = \frac{0.1 R^2}{G};$$

$$= \frac{0.1 \times (1.1 \times 10^9)^2}{6.7 \times 10^{-11}}$$

$$\approx 2.0 \times 10^{27} \text{ kg}; \quad 3$$

[12]

14. (a) change in potential energy per unit mass / work done per unit mass;
in moving small / point mass from infinity to the point;
Do not allow "from a long distance away". 2

2

- (b) (i) asymptotic at large r and in negative gravitational potential region;
line stops at surface line;
Do not allow asymptotic to y-axis. 2
- (ii) loss of gravitational potential = $6.67 \times 10^{-11} \times \frac{m_{\text{Moon}}}{r_{\text{Moon}}}$;
equates loss of gravitational potential to $\frac{1}{2}v^2$;
 $v = 2.4 \text{ kms}^{-1}$; 3
- (iii) meteorite may have initial speed / velocity towards Moon / contribution
of Earth's gravity; 1
- (c) (i) constant; 1
(ii) decreasing; 1
15. Gravitational potential
- (a) the work done per unit mass;
in bringing a small / point mass;
from infinity to the point (in the gravitational field); 3
- (b) from the graph $V_0 = 3.9 (\pm 0.2) \times 10^7 \text{ J kg}^{-1}$;
 $g_0 = \frac{V_0}{R_0} = \frac{39}{5}$;
 $= 7.8 (\pm 2) \text{ N kg}^{-1}$; 3
- (c) $2.0 \times 10^7 \text{ m}$ above surface is $2.5 \times 10^7 \text{ m}$ from centre;
 ΔV between surface and $2.5 \times 10^7 \text{ m} = (3.9 - 0.80) \times 10^7$
 $= 3.1 (\pm 0.2) \times 10^7 \text{ J kg}^{-1}$;
 $v = \sqrt{\frac{2m\Delta V}{m}} = \sqrt{2\Delta V}$;
 $= \sqrt{6.2 \times 10^7} = 7.9 (\pm 0.2) \times 10^3 \text{ ms}^{-1}$; 4
*Award [3 max] if the candidate forgets that the distances are
from the centre (answer $3.2 \times 10^3 \text{ ms}^{-1}$), ie the candidate must
show ΔV .*
- [10]
16. (a) (i) direction is changing and so there is an acceleration;
there must be a resultant force on the satellite / force is provided
by gravitational attraction; 2
- (ii) object and satellite have the same acceleration;
acceleration is towards centre of planet;
so no reaction force between object and satellite; 3

3

- (b) (i) potential energy $\frac{-(Gmm)}{(R+h)}$; 1
- (ii) in orbit, $\frac{mv^2}{(R+h)} = \frac{(Gmm)}{(R+h)^2}$ or expressed in words;
use of $E_K = \frac{1}{2}mv^2$;
 $E_K = \frac{1}{2} \frac{(Gmm)}{(R+h)}$; 3
- (c) (total energy = potential energy + kinetic energy)
total energy is $\frac{-(Gmm)}{2(R+h)}$;
as total energy is reduced, $\frac{(Gmm)}{2(R+h)}$ increases;
hence h decreases;
*Do not award if there is no reasoning or reasoning is fallacious
or misleading.*
 E_K increases and v increases; 4
- (d) friction reduces the total energy of the satellite;
causing height to decrease and speed to increase;
less height, greater frictional force;
because atmosphere denser;
frictional force causes heating effect;
as height decreases heating effect increases / heats up more;
if satellite small, sufficient heating to cause destruction;
Do not allow "heats up as height decreases". 4 max

1

4